Mars interior structure and spin state from InSight's RISE radio-science experiment

Attilio Rivoldini^{1,}, Sébastien Le Maistre^{1,2*}, Alfonso Caldiero^{1,2}, Marie Yseboodt¹, Rose-Marie Baland¹, Mikael Beuthe¹, Tim Van Hoolst^{1,3}, Véronique Dehant^{1,2}, William Folkner⁴, Dustin Buccino⁴, Daniel Kahan⁴, Jean-Charles Marty⁵, Daniele Antonangeli⁶, James Badro⁷, Mélanie Drilleau⁸, Alex Konopliv⁴, Marie-Julie Péters¹, Ana-Catalina Plesa⁹, Henri Samuel⁷, Nicola Tosi⁹, Mark Wieczorek¹⁰, Philippe Lognonné⁷, Mark Panning⁴, Suzanne Smrekar⁴, W. Bruce Banerdt⁴.

¹Royal Observatory of Belgium, Brussels, Belgium;²UCLouvain, Louvain-la-Neuve, Belgium; ³Institute of Astronomy, KU Leuven, Leuven, Belgium; ⁴Jet Propulsion Laboratory, California Institute of Technology, Pasadena, USA.; ⁶Centre national d'Études Spatiales, Toulouse, France. ⁶ IMPMC, Sorbonne Université, MNHN, CNRS, Paris, France. ⁷ Université de Paris, Institut de Physique du Globe de Paris, CNRS, Paris, France; ⁸ Institut Supérieur de l'Aéronautique et de l'Espace SUPAERO, Toulouse, France; ⁹ DLR Institute of Planetary Research, Berlin, Germany; ¹⁰ Université Côte d'Azur, Observatoire de la Côte d'Azur, CNRS, Laboratoire Lagrange, Nice, France.



The Rotation and Interior Structure Experiment



- **RISE** is together with **SEIS** and **HP3** one of the main instruments of the InSight mission lacksquare
- **RISE**: determine the rotation of Mars \bullet
 - precession
 - measure the nutation of the spin axes to detect and quantify the effect of the liquid core

• measure the rotation rate of Mars on a seasonal timescales to constrain the atmospheric angular momentum budget

RISE setup

\Rightarrow uses radio-links to reconstruct the motion of the lander in space



- **determine**: precession rate and nutation

Coherent transponder and 2 horn-antennas fixed on the Mars surface

measure: lander position, rotation rate, rotation in space



- the gravitational torque exerted by the Sun on the flattened rotating Mars causes a precession of the rotation axis in space (~171000 years)
- torque variations due to the relative positions between the Sun and Mars lead to periodic motions of the rotation axis, the nutations (1/(1,2,3,4..) year) \Rightarrow lander position changes by about 10 m on the surface

Precession and nutation



Nutation: interior structure

- inertia and from the tidal potential (well known forcing periods)
- mode, the Free Core Nutation
- if the FCN frequency ω_{FCN} is close to forcing frequency ω the nutation amplitude can be resonantly amplified

the interior structure of the planet \Rightarrow moments of inertia of the planet (A) and core (A_f), planet (e) and core shape (e_f), core compliances due to tidal forcing (γ) and rotation rate variation (β)

• if a planet were rigid then nutation amplitudes can be predicted very precisely from its moment of

• nutation amplitudes depend on the interior structure of Mars and in particular on the liquid core

• the relative rotation between the fluid core and solid mantle is characterised by a rotational normal



Liquid core signature and real data





- the measured doppler shift is about 20-30 mHz
- the signature of the liquid core is 2 orders of magnitude smaller
- since its periods are well known and because of data accumulation it can be determined





Time after landing (Earth days)

- amplification factor F and FCN period $\tau_{\rm FCN}$
- $F = 0.061 \pm 0.0064$ and $\tau_{\text{FCN}} = -242.25 \pm 2.7$ days
- F in expected range but $\tau_{\rm FCN}$ somewhat lower than expected

RISE results

>600 days of data are required to obtain robust estimates for the core

Prior modelling assumption

- layered spherical isotropic structure (crust, mantle, core)
- 1999 (**EH45**))
 - properties

 - \bullet al. 2018)



• bulk mantle chemical composition deduced from Martian meteorites and assumptions about its formation matching either refractory elements (i.e. Wänke & Dreibus 1994, Taylor 2013, Yoshizaki & McDonough, 2020 (YMD)) Or Oxygen isotope systematics (i.e. Lodders & Fegley, 1997; Sanloup et al.,

• mantle rich in FeO (~18wt%) except YMD model (~14wt%) \Rightarrow affects depth of major phase transition, density, and elastic

• mantle is chemical homogeneous and in thermodynamical equilibrium \Rightarrow use Gibbs energy minimisation (Perple_X (Connoly 2005) + Stixrude & Lithgow-Bertelloni 2011) to compute all thermoelastic properties in the mantle from bulk composition and thermal state \Rightarrow density, bulk moduli, and shear moduli are related by bulk composition and remaining degree of freedom: thermal state

assume a hot and cold mantle temperature end-member deduced from 3D spherical geometry thermal evolution studies (Plesa et

- core is fluid, homogeneous, and isentropic
 - candidate light elements that are siderophile or dissolve into Fe at formation and affect thermoelastic properties of the core: S, O, C, H (e.g. Steenstra 2018)
 - bulk composition models and formation allow for **S**≲**17wt%** (e.g. Steenstra 2018)
 - **O** (≤4wt%) fraction dependents on mantle FeO and bulk S (Gendre & Badro 2022)
 - maximal amount of C (≤1.5wt%) limited by S (Dasgupta 2016)
 - **H** (<**1wt%**) depends on initial mantle H₂O and extremely scarce experimental data (Tagawa 2022)

et al. (1999); Taylor (2013) iizaki and McDonough (2020 $_{\rm ntle}^{\rm ntle} = 14.7 \ {\rm wt.} \ \%$ Oxygen fraction (wt. Co<mark>smochemi</mark>cal constraints 25 15 20 30 Sulfur fraction (wt. %)





- mantle boundary) (see also Zharkov 2009, Wieczorek 2019)
- internal loads are specified to match geometric and dynamic shape of Mars
- effect of rotational flattening on core shape ~5000 m (also dependent on density jump at the CMB)
- effect of internal loads on core shape ~100m

geometric and dynamic shape (deg 2) of Mars results from rotation, mass anomaly induced by the surface topography, and internal mass anomalies placed deep within the planet (Moho, bottom of the lithosphere, core-

Interior structure before InSight

- constraints mostly from geodesy data (gravity field, precession rate, tides)
- observations (but often disagree on core composition)
- measured tides can only be explained with a liquid core (Yoder 2003)
- cessation of core generated dynamo about 3.7 Ga ago (Mittelholz 2020) \rightarrow inner core unlikely?
- core radius 1790±65km (e.g. Rivoldini 2011, Khan 2019)
- inferred core density can be explained with a liquid Fe-16.5wt%S alloy molten
- Terasaki 2020, Xu 2021, Nishida 2020)
- \rightarrow if S were the sole light element would lead to up-floating solid FeS

• structure models based on mantle compositions deduced from bulk composition models agree with

 \rightarrow about the eutectic concentration any plausible present-day thermal state implies that inner core is fully

• new experimental data about elastic properties (density and acoustic velocity) of liquid Fe-S show that more than 25wt%S is required to match the density of the core (Nishida 2016, Shimoyama 2016, Kawaguchi 2017, Morard 2018,



Precession: interior structure

- lacksquare
- with the data!



• precession rate together with degree 2 gravity field provides polar moment of inertia

leads to constraints on the mass distribution within Mars and in particular on the crust

• allows to further constrain crustal models deduced from seismic data, surface gravity, and topography → Wänke & Dreibus 1994 and Taylor 2013 BSM models incompatible

Nutation: interior structure



- well as nearly independent on core shape
- and density profile affect hydrostatic core shape
- (Stähler et al. 2021, Duràn et al 2022)

• F (~core moment of inertia~ core radius) weakly dependent on mantle composition and temperature as

but **not** the FCN period: precession compatible crust density and thicknesses affect internal loading

inferred core radius range of 1825±55 km is in excellent agreement with tides and seismic observations

Nutation: Core composition

- RISE data and geochemical constraints require a core with 17±2.5wt% S, agreement with arrival times of core traversing seismic waves (Irving 2023) 2014, Shimoyama 2016, Terasaki 2010, Kawaguchi 2017, Tagawa 2022)



models without H are unlikely if S in agreement with geochemical constraints

2.5±0.5wt% O, and 1.0±0.5wt% C if 0.75wt% H is assumed in the core in good

(eos used for core composition deduction: Dorogokupets 2017; Nishida 2016,2020; Morard 2017, 2018; Xu 2021; Komabayashi

What about the inner-core?

- FCN lead to a resonant amplification of the nutation
- the signature of the inner core has not been detected in RISE data small and its effect on nutation is below current precision

- very early cessation of the core generated dynamo (~3.7 Ga ago) in favour of a purely thermally driven dynamo

• if there is an inner core it can have a relative rotation with respect to the fluid core giving rise to a another rotational normal mode the Free Inner Core Nutation (FICN) that can like the

 \rightarrow either the FICN period differs significantly from tidal forcing periods or the inner is rather

• core traversing seismic waves (SKS) are not reflected inside the liquid core and sample the core down to a radius of about 800km (Irving 2023) \rightarrow the inner-core is smaller than 800km

• an S concentration close to the Fe-S eutectic together with the naïve assumption that the remaining light elements further reduce the melting temperature of Fe requires an implausible low core temperature (<1300K) to initiate solidification in the core \rightarrow the core is fully molten

Conclusions

- and the shape of the core
- seismic data
- 1.0±0.5wt% C if 0.75wt% H is assumed
- a small fraction of H is required to obtain an acceptable fraction of S
- anomalies

• the measured nutation and the detection of the FCN normal mode confirm the liquid state of the core

RISE data constrain the moment of inertia of the core, the density jump at the core mantle boundary,

• the core radius of 1825±55 km is in excellent agreement with estimates obtained from tides and

• RISE data and geochemical constraints require a core with 17±2.5wt% S, 2.5±0.5wt% O, and

• the FCN period can be explained if the core has an almost hydrostatic shape, such a core shape can result from deep seated mass anomalies within the mantle that originate form thermal or chemical

Caveats and motivation for future work

- liquid Fe-O-S-C-H alloys are unknown
- eos of liquid Fe-H at Mars core conditions is really not well known (almost not)
- stratified (e.g. Fe-O-S (Tsuno 2007), Fe-S-H (Yokoo 2022)

- core is smaller and denser?
- this explain the LVZ like layer in the lower evidenced in seismic data? (Durán 2022)

• thermal evolution and thermal state of the core cannot are difficult to assess since transport properties of

• current eos of the core based on ideal mixing assumption of Fe-S, Fe-O, Fe-C, and Fe-H but it is well known that the individual light elements do not mix ideally with $Fe \rightarrow$ bias inferred core composition

• non-deal mixing will lead to immiscibility fields in the composition space \rightarrow the core could be chemically

• joint effect of O,S,C, and H on melting temperature of Fe unknown \rightarrow is the liquidus really that low?

• thermodynamic database of mantle minerals calibrated on samples close to earth mantle composition

• suggested presence of present-day molten layer in the bottom of the mantle (Samuel 2021) \rightarrow is the metallic

• LVZ and soft lower mantle due to the interaction of liquid Fe-S with olivine aggregates (Kono 2023) \rightarrow can

Interpretation: FCN period

- FCN frequency proportional to core shape which is directly related to the density jump at the core mantle boundary (⇒ constraints density jump at the core-mantle boundary)
- RISE data implies an almost hydrostatic core shape, but the shape of Mars is not hydrostatic
 - requires mass anomaly at the bottom of a thick lithosphere (>550km)
 - or two loads at shallow depth and at the core-mantle boundary

